



# The NO $\nu$ A Experiment

(NuMI Off-Axis  $\nu_e$  Appearance)

Robert Bernstein

Fermilab

NO-VE 2006

Neutrino Oscillations in Venice

7 February 2006

(For the Collaboration)



# Outline for Talk

- Physics Goals:
  - A developing, staged program:
    - $\theta_{13}$  ( $\nu_e$  Appearance)
    - Sign of  $\Delta m^2$
    - CP-Violation
- The NOvA detector:
  - Size, structure, resolution
  - Comparison to MINOS
  - Near detector
- Simulations and Backgrounds:
  - What is done and what is underway
- Status, Cost & Schedule



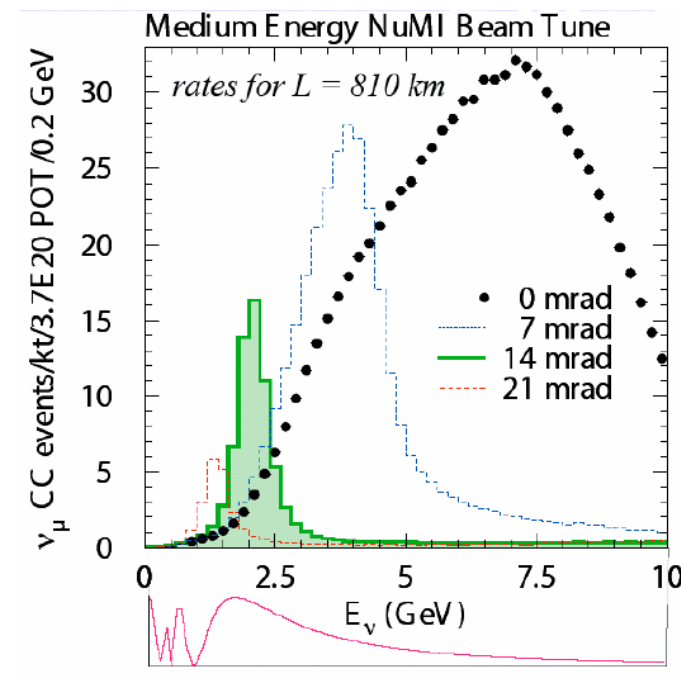
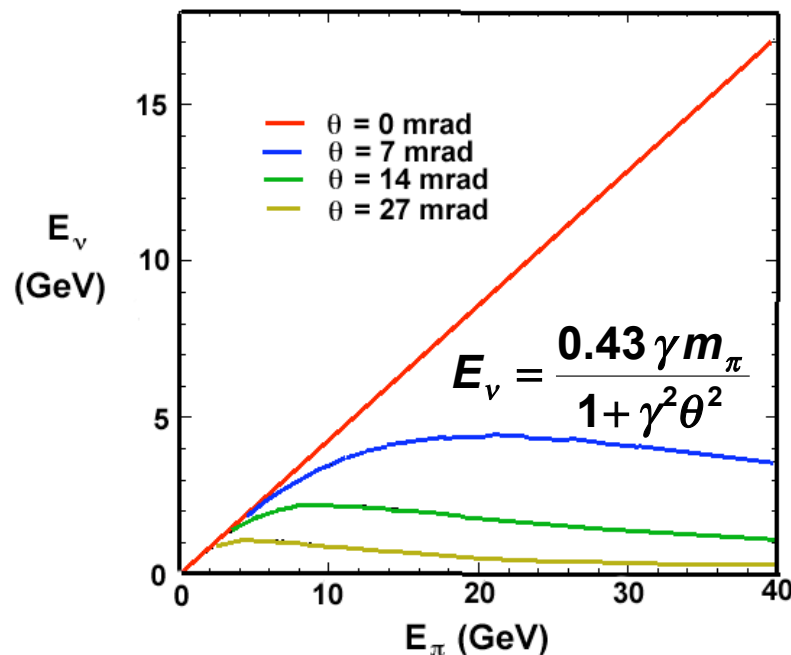
## Current Ideas & Goals

- Phase 1 (Extend  $\nu_\mu \rightarrow \nu_e$  reach by order of magnitude; explore mass hierarchy & CPV)
  - NOvA
  - T2K (Kajita and Nishikawa)
  - Reactors:  $\theta_{13}$  (Laserre)
- Phase 2 (Extend reach for mass hierarchy & CPV)
  - NOvA + PD (Geer)
  - HK + PD
  - Brookhaven VLBL proposal
- CERN PD + Mton H<sub>2</sub>O detector
- Phase 3 (Significantly extend reach)
  - Neutrino factory
  - Beta beam



# NuMI Off-Axis

- 15 mrad off axis & 810 km from the source
- Yields a narrow band beam on oscillation maximum
- More flux & reduced backgrounds from the HE tails

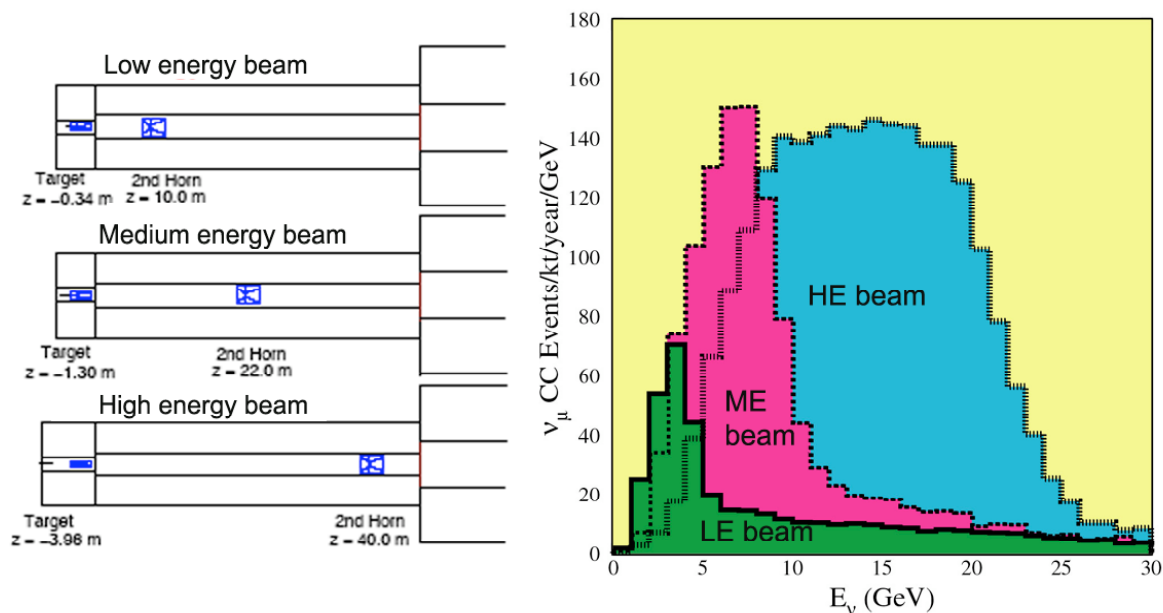






# NuMI Beam

- Move Horn toward/away from target:



cross-section included



# Rates and Spectra

- Spectra for NuMI Medium Energy Beam:

Source of  $\pi^0$ 's to be cut

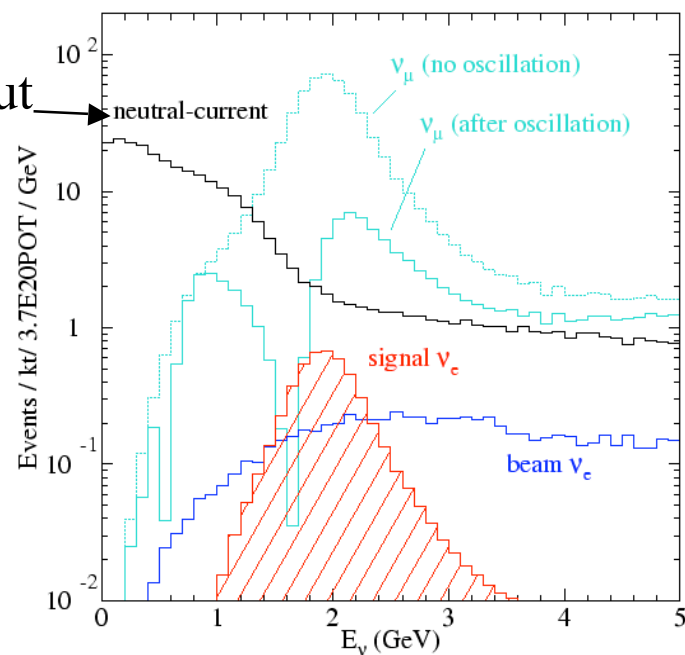


Fig. 2.8: Simulated energy distributions for the  $\nu_e$  oscillation signal, intrinsic beam  $\nu_e$  events, neutral-current events and  $\nu_\mu$  charged-current events with and without oscillations. The simulation used  $\Delta m^2_{32} = 2.5 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2(2\theta_{23}) = 1.0$ , and  $\sin^2(2\theta_{13}) = 0.04$ . An off-axis distance of 12 km at 810 km was assumed.



## $P(\nu_\mu \rightarrow \nu_e)$ in Vacuum

$$P(\nu_\mu \rightarrow \nu_e) = P_1 + P_2 + P_3 + P_4$$

$$P_1 = \sin^2(\theta_{23}) \sin^2(2\theta_{13}) \sin^2(\Delta m_{13}^2 \frac{L}{E}) \quad \text{Atmospheric}$$

$$P_2 = \cos^2(\theta_{23}) \sin^2(2\theta_{12}) \sin^2(\Delta m_{12}^2 \frac{L}{E}) \quad \text{Solar}$$

$$P_3 = \mp J \sin(\delta) \sin(\Delta m_{13}^2 \frac{L}{E})$$

$$P_4 = J \cos(\delta) \cos(\Delta m_{13}^2 \frac{L}{E}) \quad \begin{array}{l} \text{Atmospheric-} \\ \text{solar interference} \end{array}$$

where

$$J = \sin(2\theta_{13}) \cos(\theta_{13}) \sin(2\theta_{12}) \sin(2\theta_{23}) \\ \times \sin(\Delta m_{13}^2 \frac{L}{E}) \sin(\Delta m_{12}^2 \frac{L}{E})$$

- note terms with  $\Delta m_{13}$  and  $\delta$  mix, hence matter-CP ambiguity



## $P(\nu_\mu \rightarrow \nu_e)$ (in Matter)

- In matter **at oscillation maximum**,  $P_1$  will be approximately multiplied by  $(1 \pm 2E/E_R)$  and  $P_3$  and  $P_4$  will be approximately multiplied by  $(1 \pm E/E_R)$ , where the top sign is for neutrinos with normal mass hierarchy and antineutrinos with inverted mass hierarchy

$$E_R = \frac{\Delta m_{13}^2}{2\sqrt{2}G_F\rho} \approx 11 \text{ GeV for the Earth's crust}$$

- About  $\pm 30\%$  effect for NuMI/NovA
- About  $\pm 11\%$  effect for T2K
- The effect is reduced for energies above the oscillation maximum & increased for energies below



# Ambiguity between Mass Ordering and CP Phase

- Need  $\nu$  and  $\bar{\nu}$
- Ambiguity depending on sign of  $\Delta m^2$ ,  $\delta$  CP

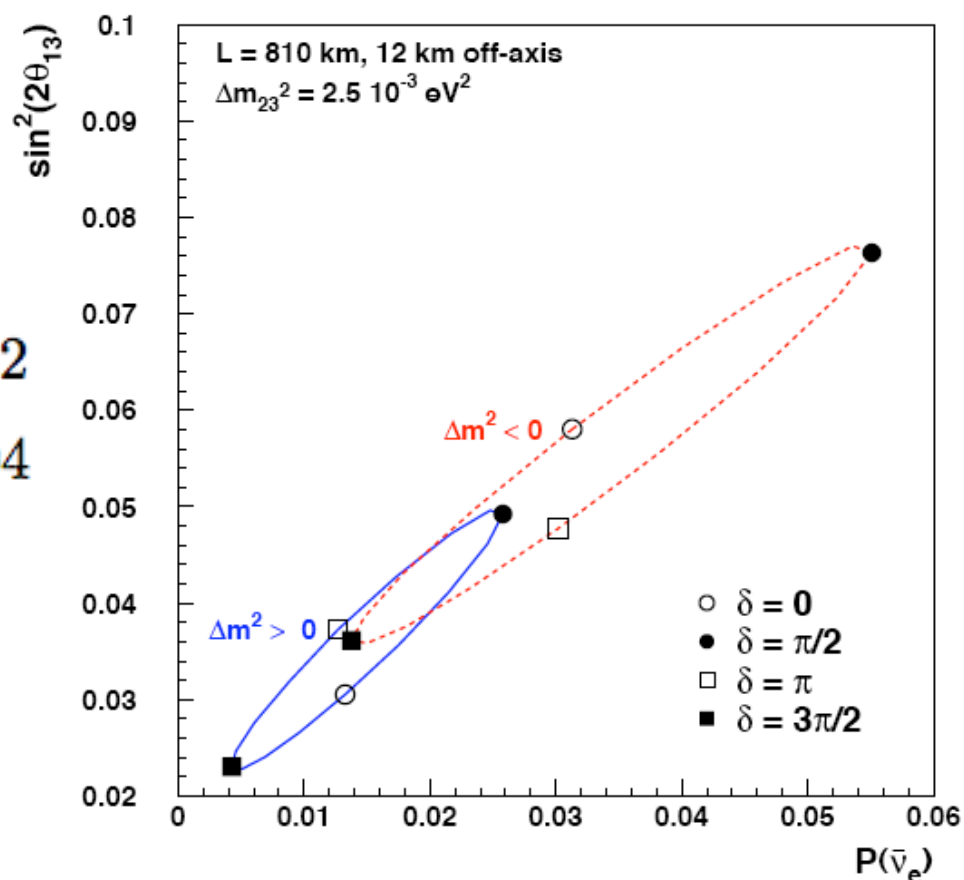
- Statistical Errors

$$P(\bar{\nu}_e) = 0.02 \pm 0.002$$

$$\sin^2 \theta_{13}(\nu) = 0.05 \pm 0.004$$

- More detail later
- Combining Nova and T2K also useful!

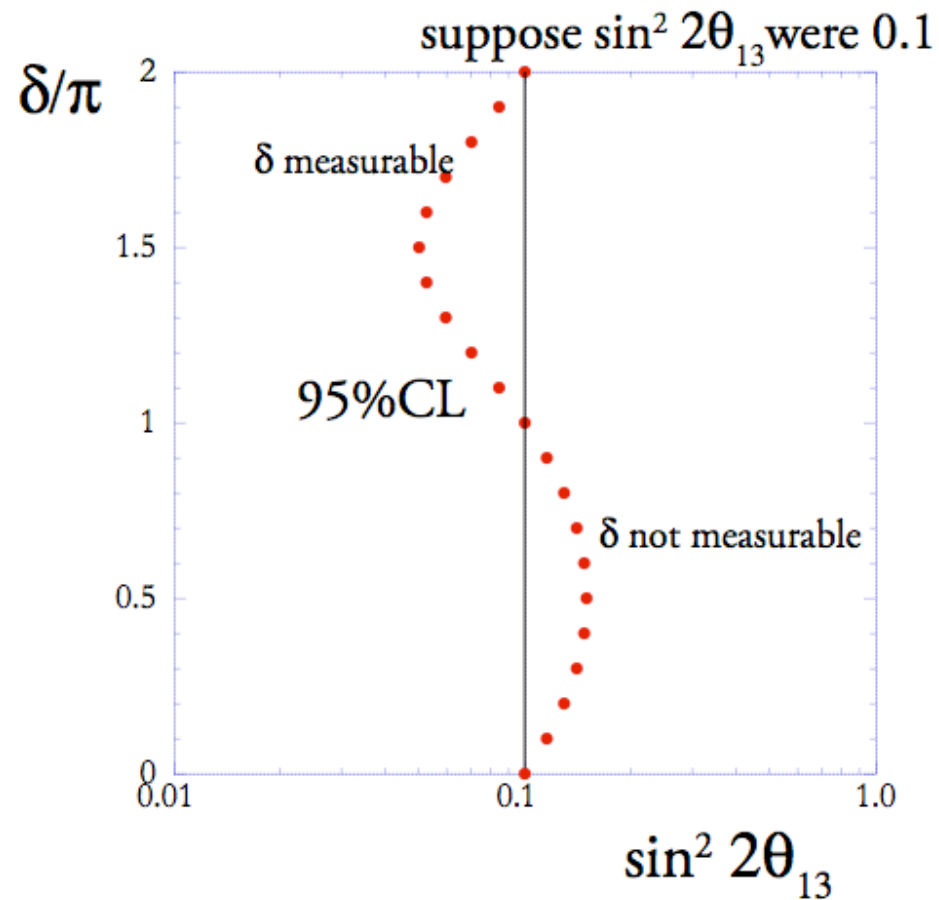
$\sin^2(2\theta_{13})$  vs.  $P(\bar{\nu}_e)$  for  $P(\nu_e) = 0.02$





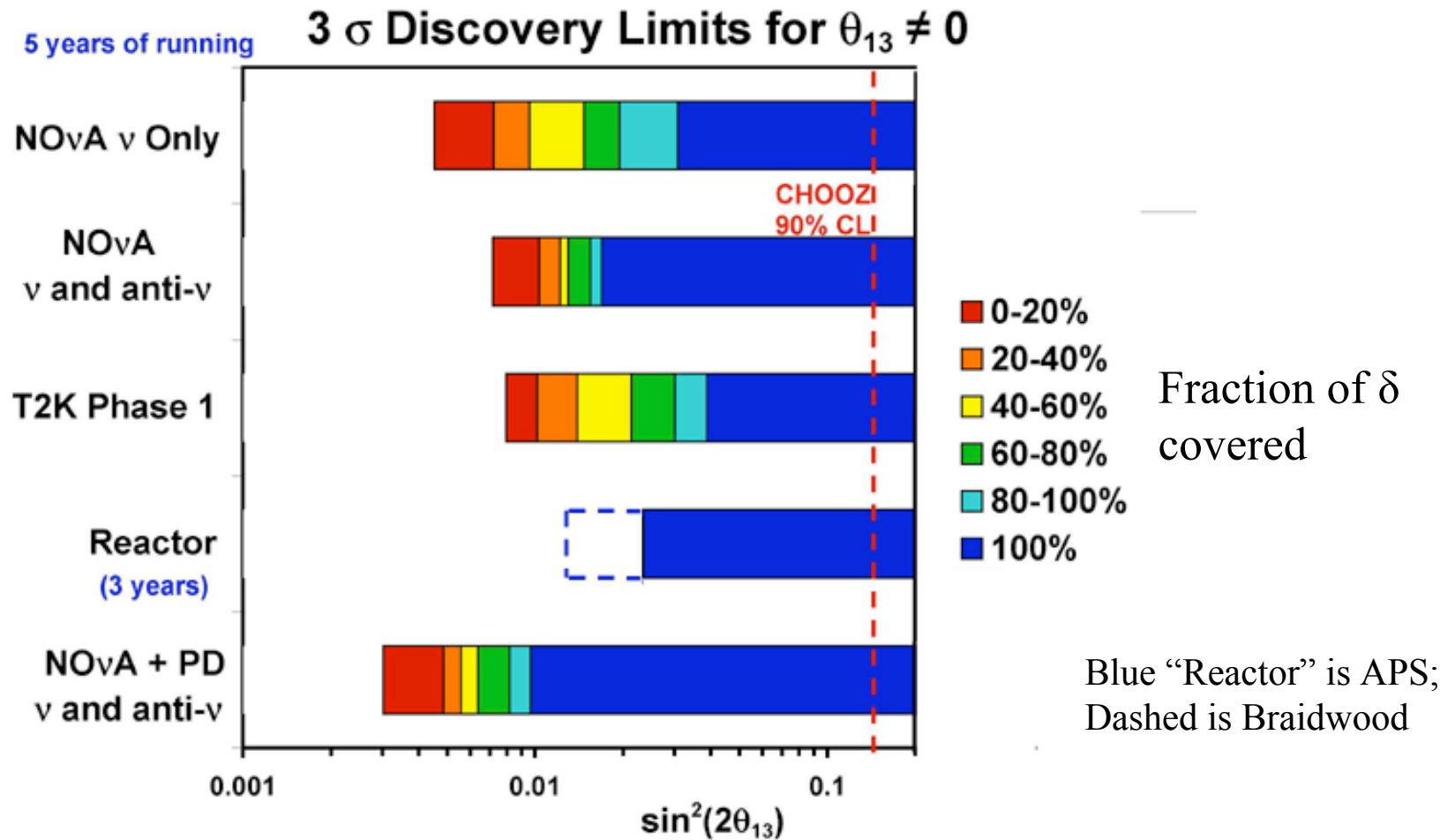
# “Fraction of $\delta$ ”

Fraction of  $\delta$ :  
Above or below  
significance  
cutoff?





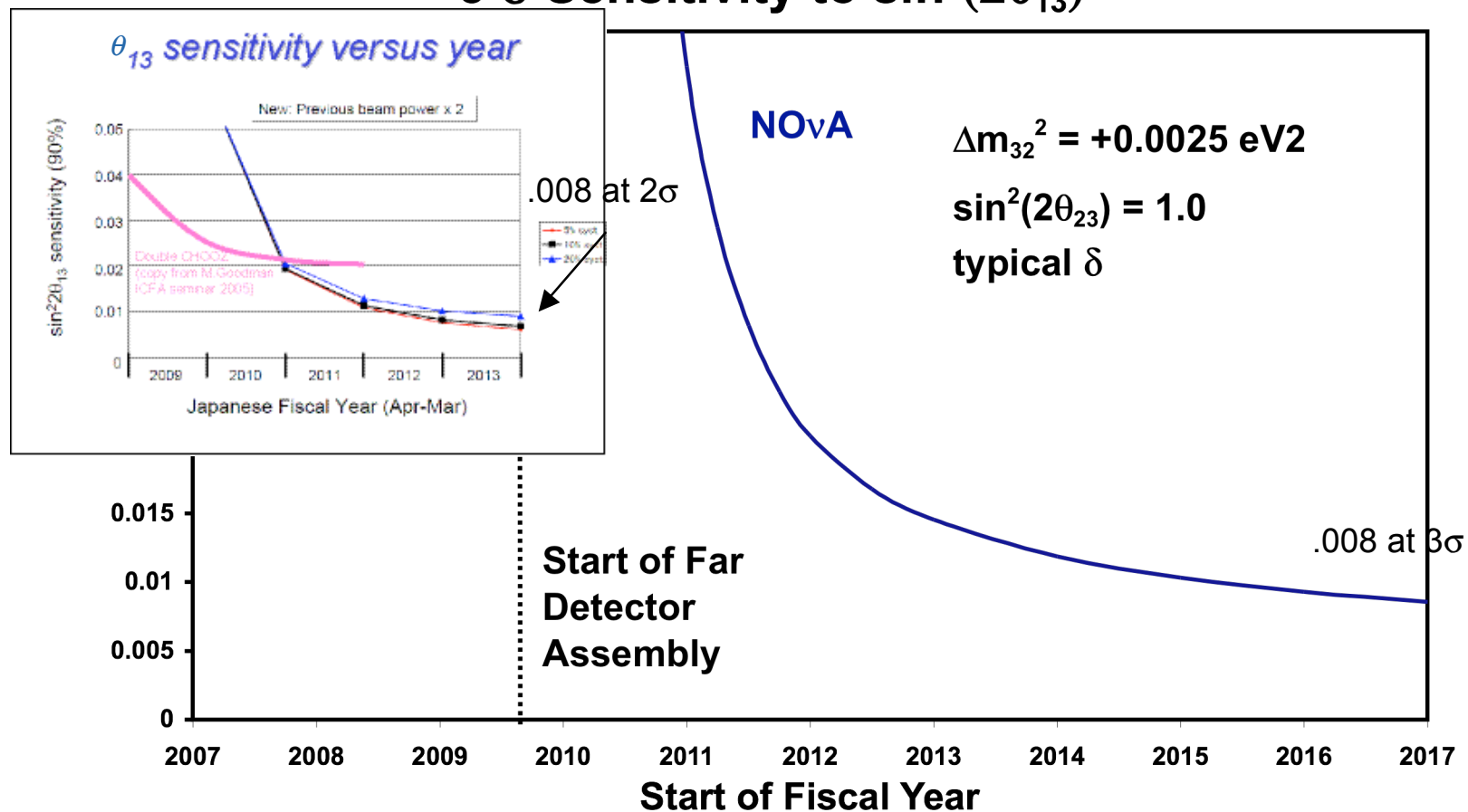
# 3 $\sigma$ Sensitivity to $\theta_{13} \neq 0$





# Sensitivity vs Time

## 3 $\sigma$ Sensitivity to $\sin^2(2\theta_{13})$







# Importance of the Mass Ordering

- Window on very high energy scales
  - Grand unified theories (mostly) favor the normal mass ordering but other approaches favor the inverted ordering
- The next generation of neutrinoless double beta decay experiment care about the mass ordering
  - If inverted ordering, can decide if the neutrino is its own antiparticle
  - If normal ordering, a negative result is inconclusive
- For CPV measurement we need to resolve the mass ordering
  - Matter effects contributes an apparent CP violation that must be corrected



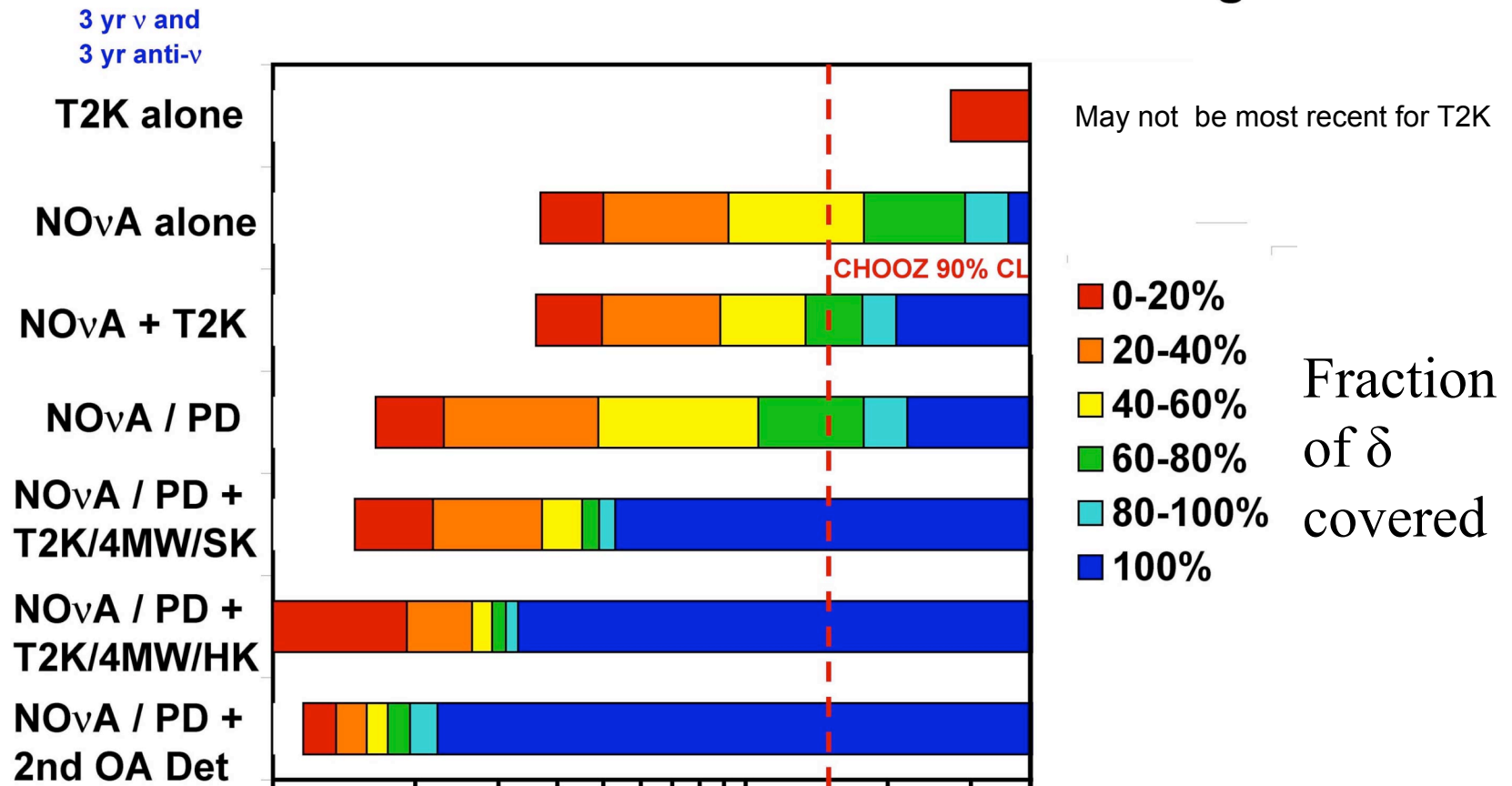
# Role of NOvA in Resolving the Mass Ordering

- The mass ordering can be resolved only by matter effects in the earth over long baselines
- NOvA is the only proposed experiment with a sufficiently long baseline to resolve the mass ordering
- The siting of NOvA is optimized for this measurement
- NOvA is the first step in a step-by-step program that can resolve the mass ordering in the region accessible to conventional neutrino beams



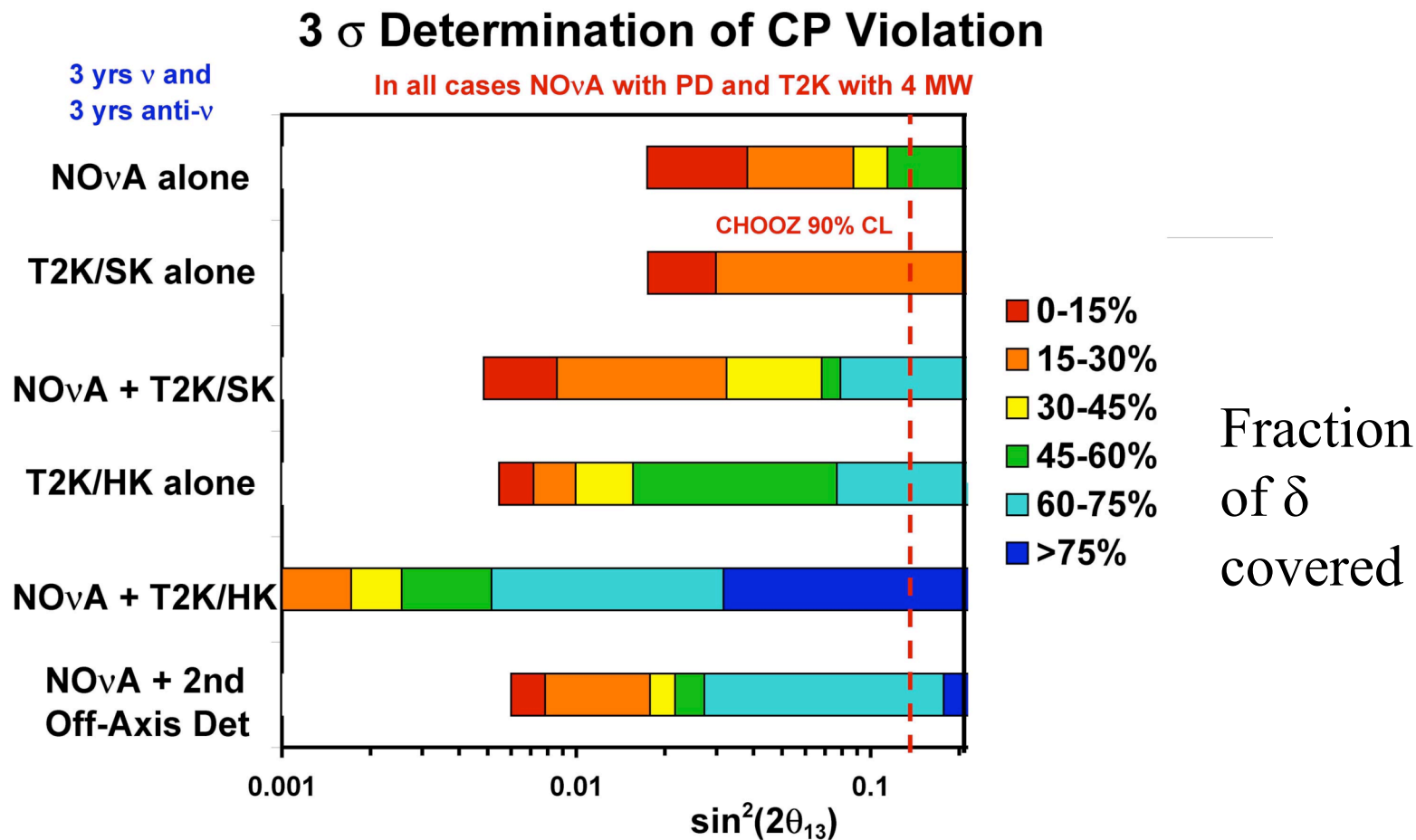
# 95% CL Resolution of the Mass Ordering

## 95% CL Determination of the Mass Ordering





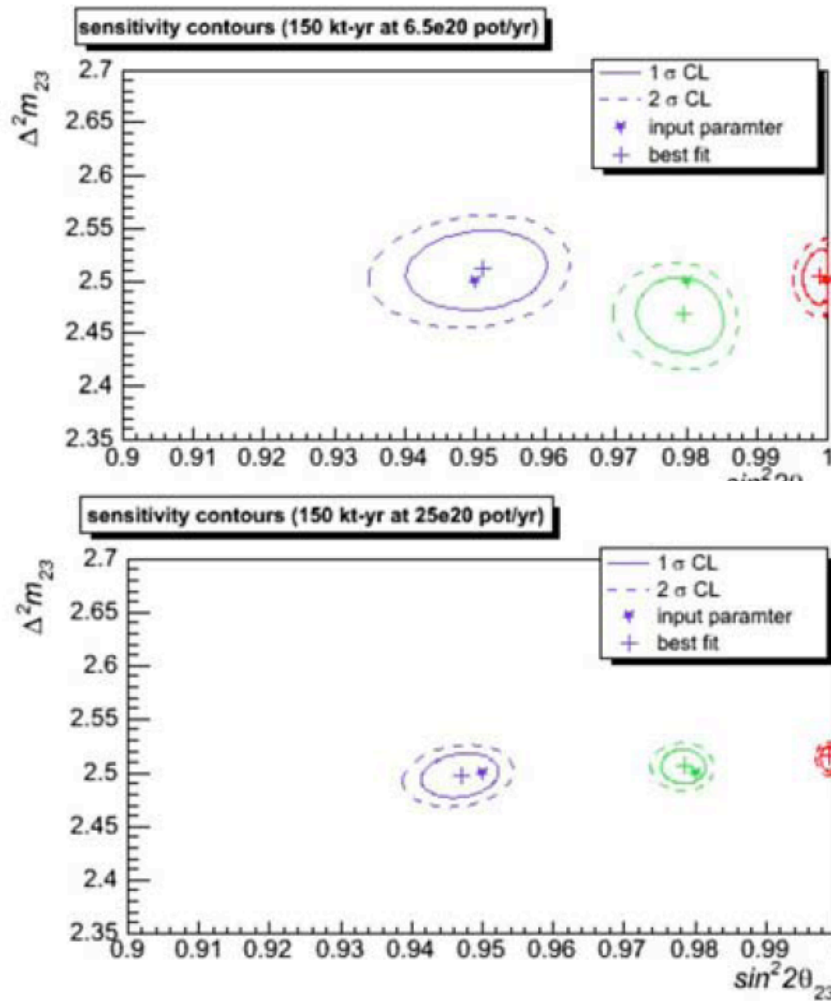
# 3 $\sigma$ Determination of CP Violation



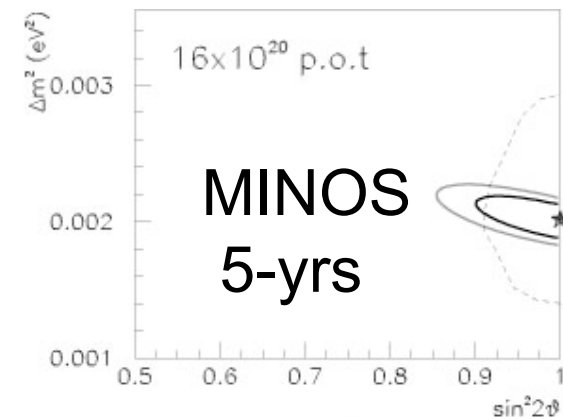
1. Must have measurement errors  $< 30^\circ$  ( $3\sigma=90^\circ$ )
2. Some values of  $\delta$  (e.g.  $\delta=0$ ) not resolvable



# Measurement of $\Delta m_{32}^2$ and $\sin^2(2\theta_{23})$



**5-year ν run**



**5-year ν run  
with Proton Driver**



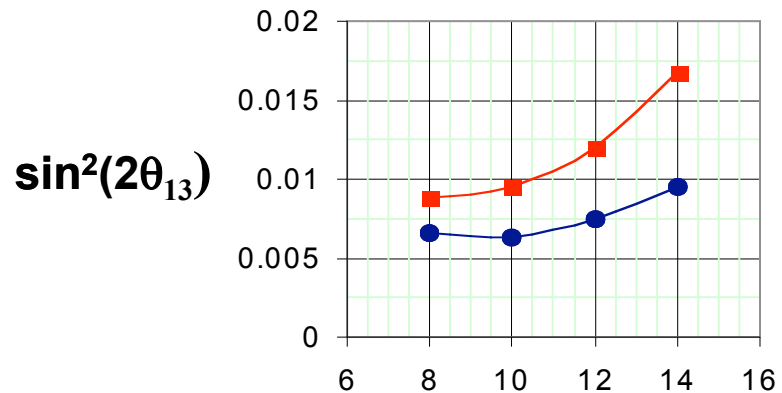
# Sensitivity to Off-Axis Distance

- 3 standard deviation discovery limit of NOvA for  $\nu_\mu \rightarrow \nu_e$  oscillations as a function of  $\sin^2(2\theta_{13})$  and the off-axis distance.
- 95% confidence level for NOvA to resolve the mass hierarchy as a function of  $\sin^2(2\theta_{13})$  and the off-axis distance.

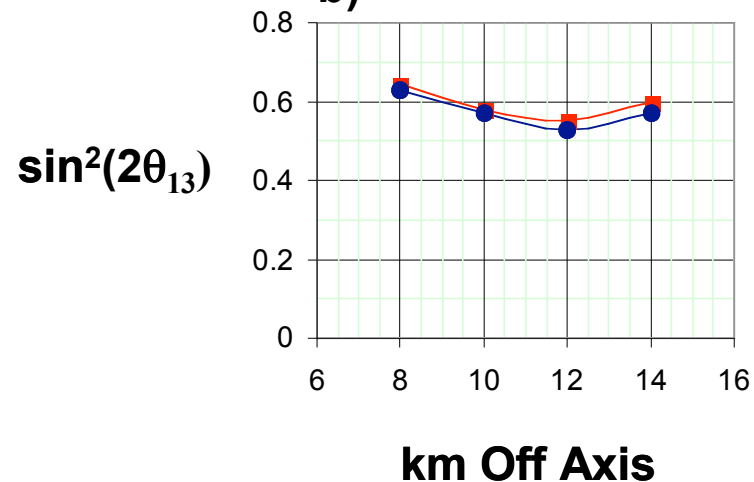
(new info from Parke?)

red normal, blue inverted

a)



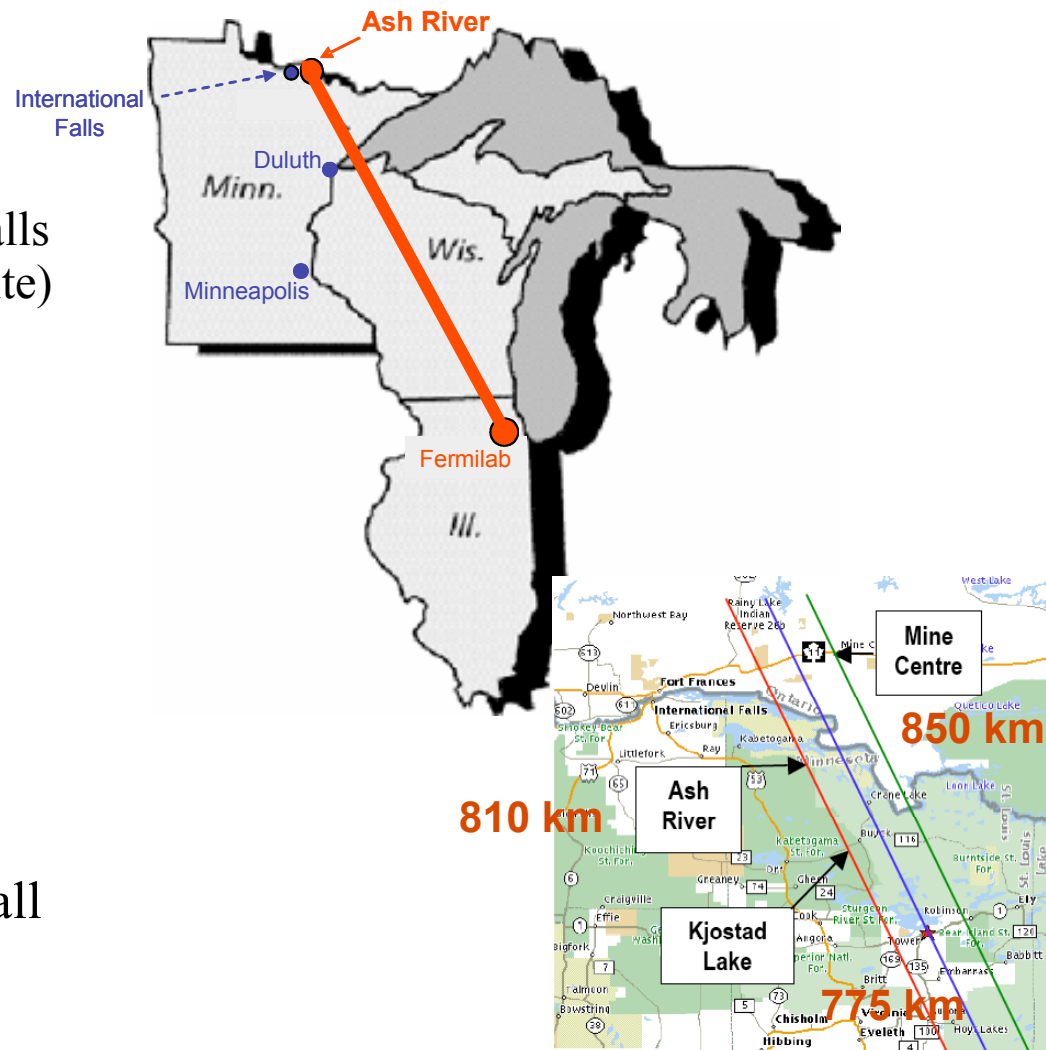
b)





# Far Detector Site

- Ash River
  - Baseline site
  - Can fly into International Falls (less than an hour drive to site)
  - Resort area
- Kjostad Lake
  - Orr-Buyck Rd
  - Good backup site
- Mine Centre
  - Must cross the US-Canada border (in a paper mill site) to get from airport to site
- Range of angles available at all sites

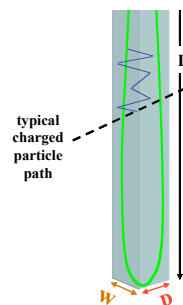
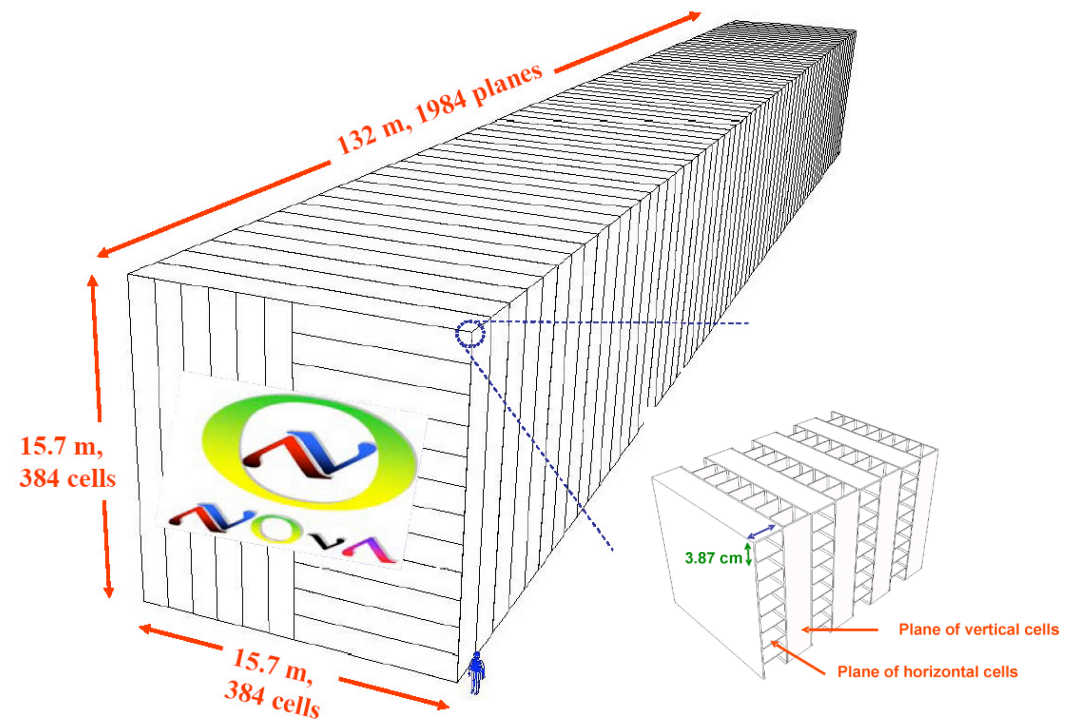






# NOvA Far Detector

- Revised Proposal (April 05)
  - “Totally Active” scintillator detector (75%)
- 30 kT
  - 24 kT liquid scintillator
  - 6 kT of cells
- Liquid scintillator cells
  - 3.9 cm x 6 cm x 15.7m
  - 0.15  $X_0$  sampling
  - 1984 planes of cells
- Cell walls
  - Extruded rigid PVC
  - 3 mm outer; 2 mm inner horizontal; vertical 4.5 mm outside
- Readout
  - U-shaped 0.8 mm WLS fiber : 27,000 km!
  - APDs (80% QE)



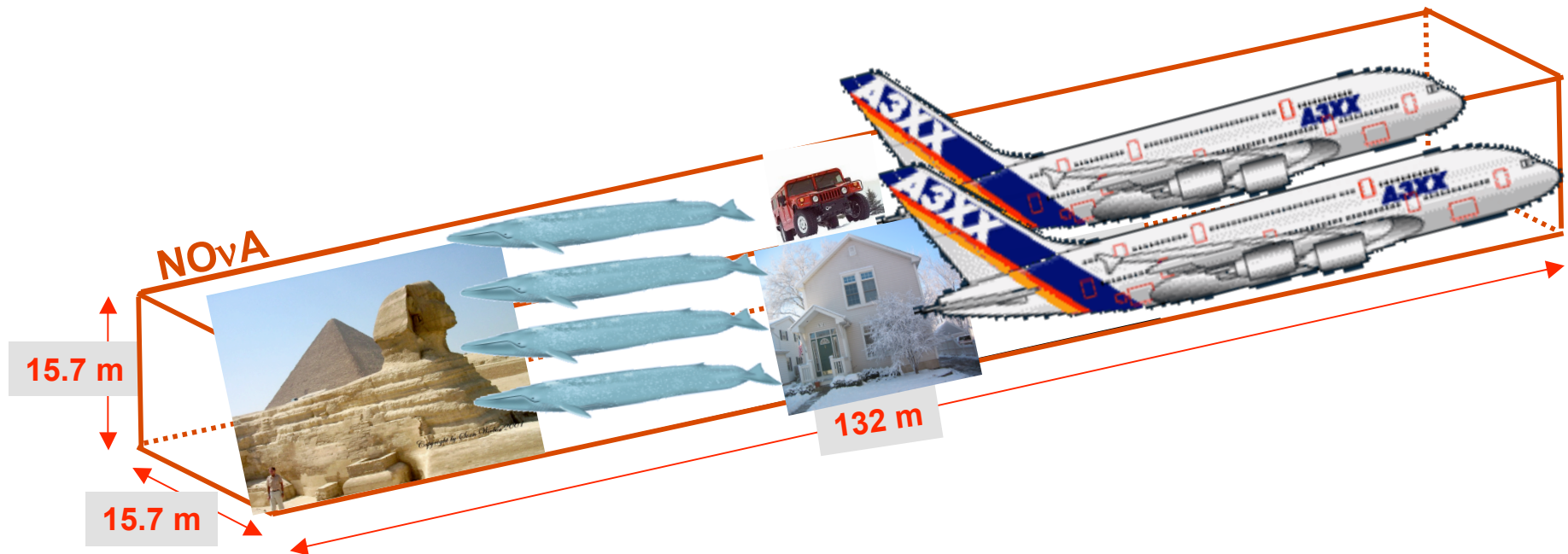
*One Plane: 15K kg*

Horizontal/vertical cells  
different for structural strength





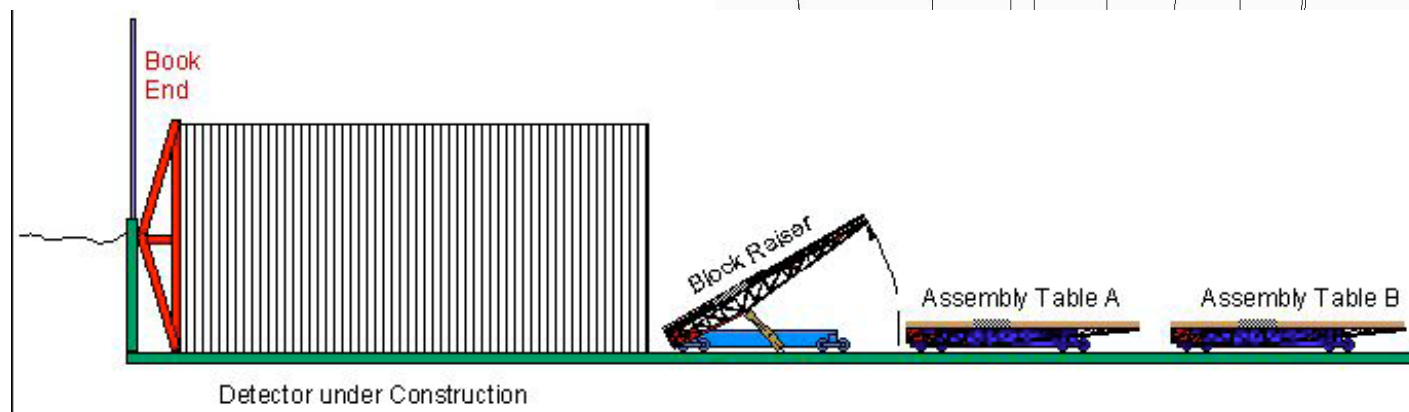
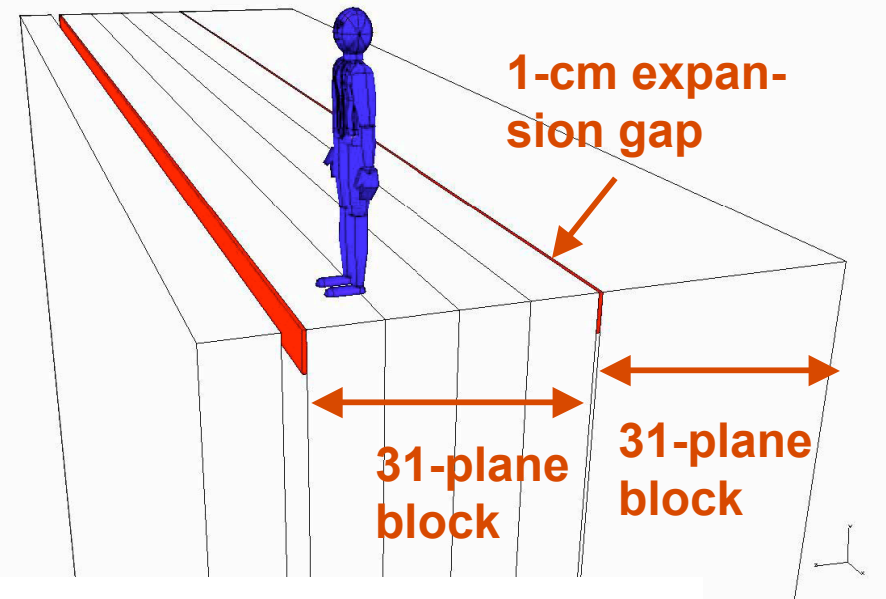
# Just how big *is* That?





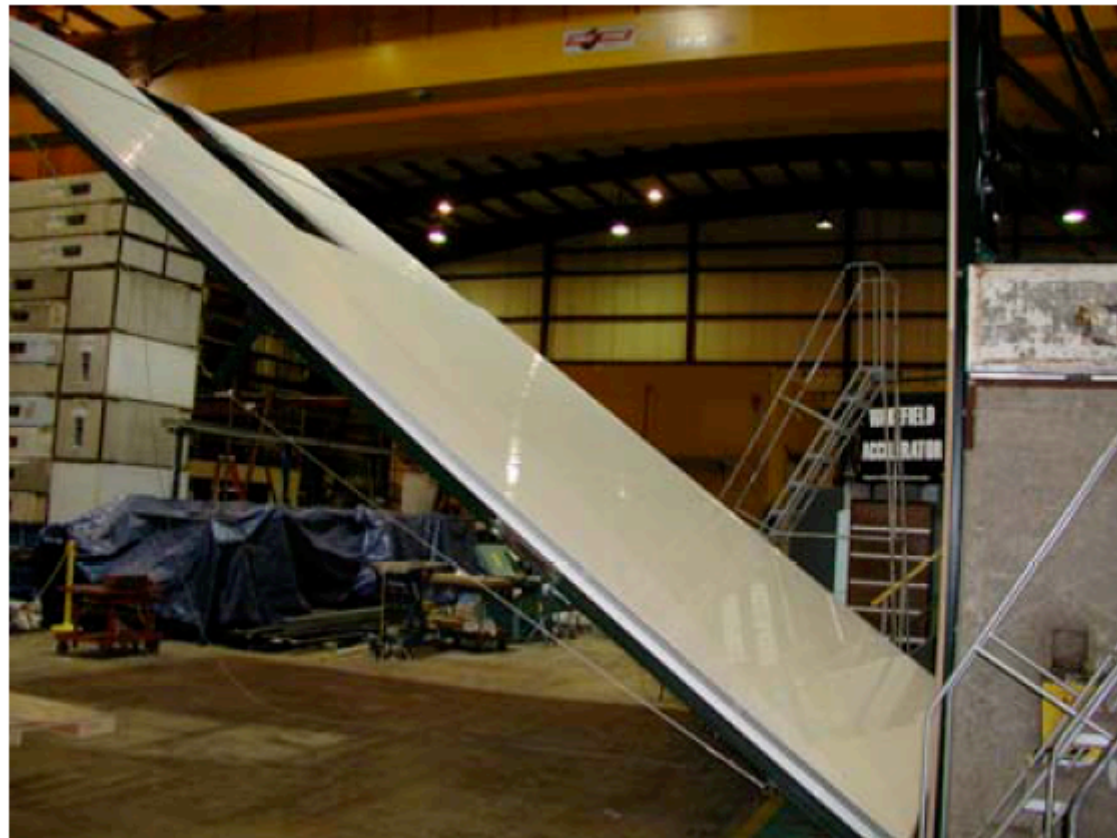
# Far Detector Assembly

- Detector has 64 (31-plane) blocks
- Can fill with scintillator and run during construction
- Half-Size planes built & tested at Argonne





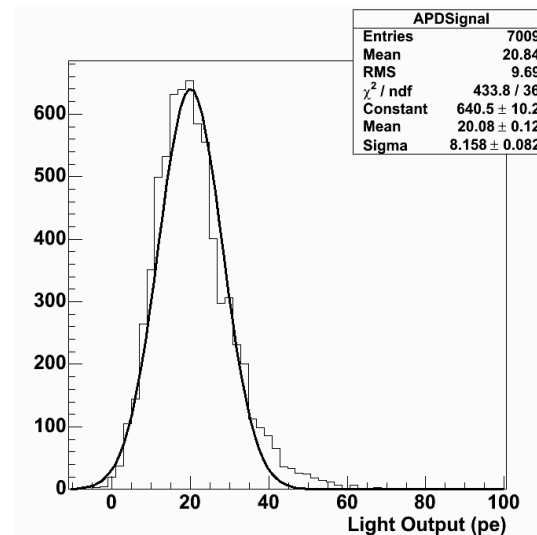
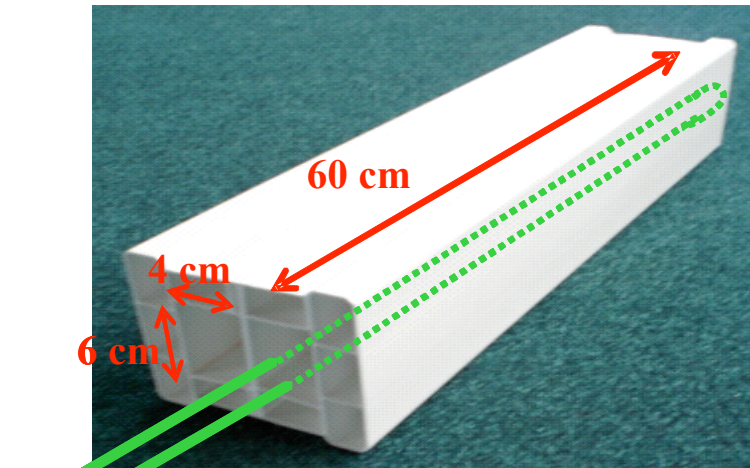
## Half Block Prototype Being Built





# Number of Photoelectrons

- Transverse Cell  
Dimensions the same
- Length of cell shorter but 33.5 meter fiber identical
- Read out to APD
- #PE stable over 2 months
- *Enough light!*





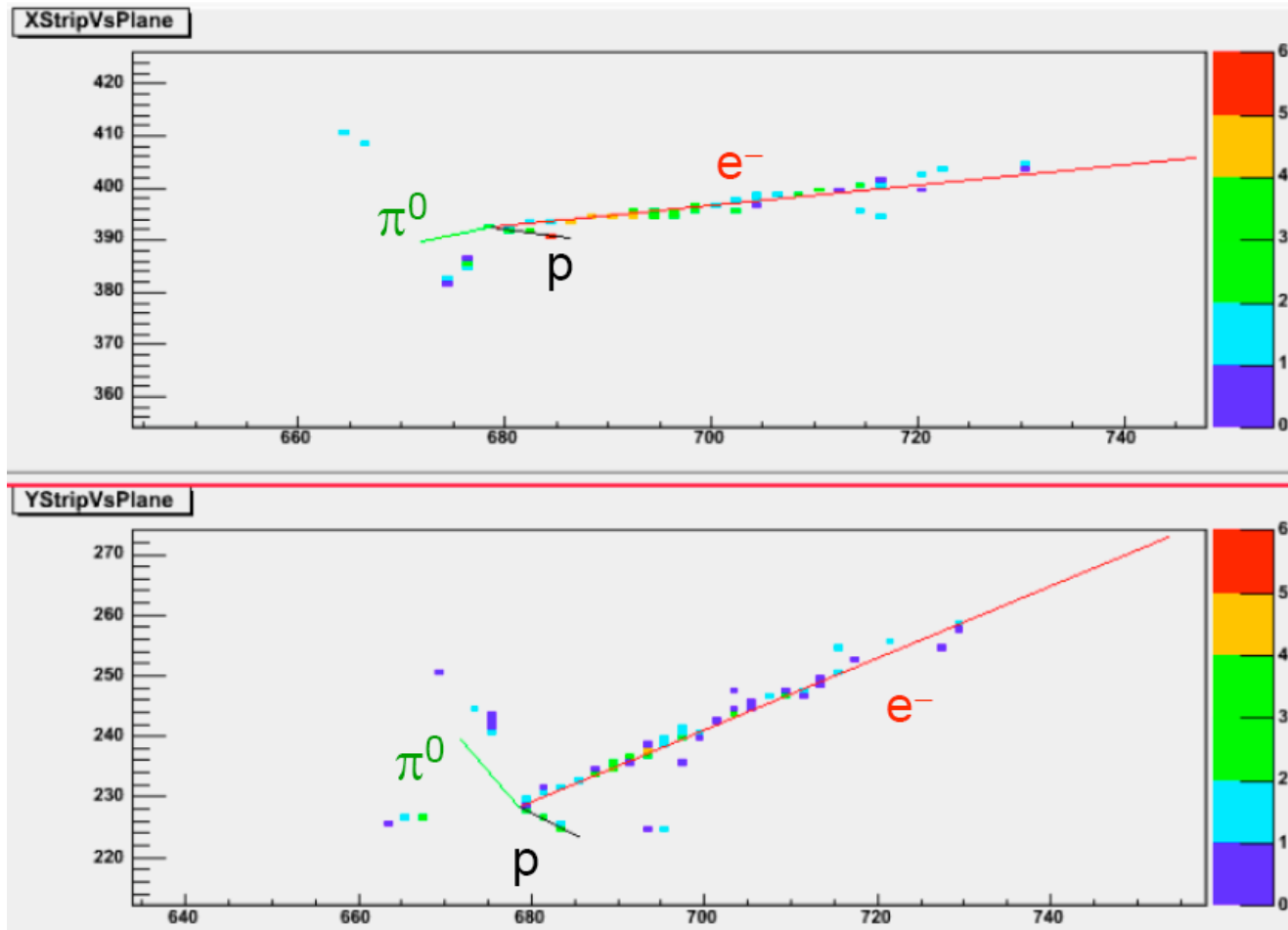
## NOvA vs. MINOS for $\nu_e$

- Other than the off-axis beam...
  - Mass increased by a factor of 6
  - Improved energy resolution and  $e/\pi^0$  discrimination
    - “Totally active scintillator detector” (TASD) instead of sampling calorimeter
    - Increase longitudinal sampling by a factor of 10  
( $1.5 X_0$  to  $0.15 X_0$ )
- All while reducing the cost/kiloton by a factor of 3 with respect to MINOS Far Detector
- No magnetic field in NOvA



# Sample MC Event

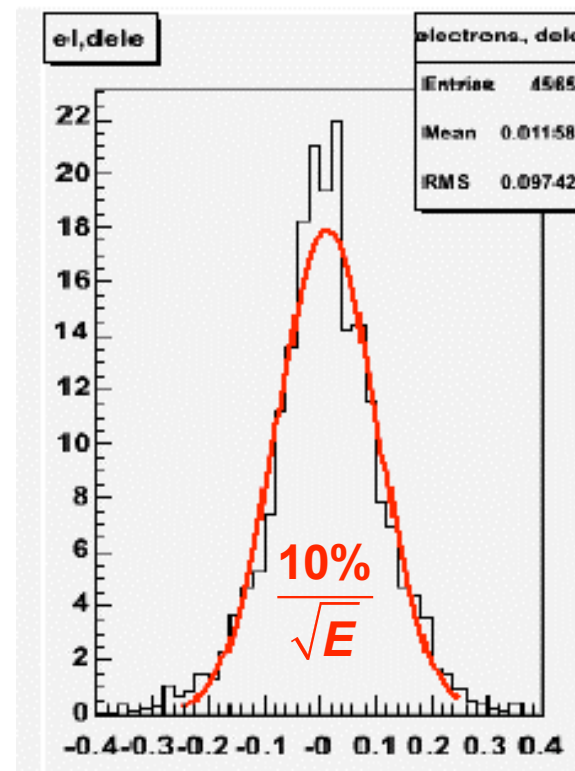
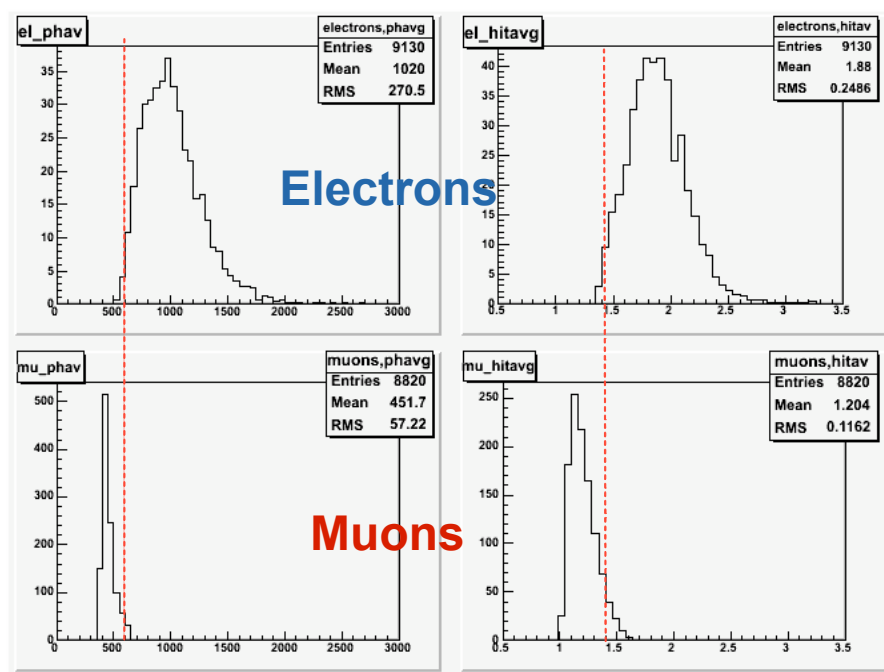
$1.65 \text{ GeV } \nu_e N \rightarrow e p \pi^0$







# Electron ID & Energy Resolution



Also use RMS of pulse height per plane,  
gaps & energy cuts



# Reconstruction

Two Algorithms: using a maximum likelihood analysis with the following variables

- a) Total measured energy
- b) Fraction of total energy carried by the electron
- c) Mean pulse height near the origin of the electron
- d) Pulse height per plane for the electron
- e) Number of hits per plane for the electron
- f) Energy upstream of the vertex
- g) Curvature of the electron
- h) Missing transverse momentum
- i) Fraction of total electron energy contained in the first half of the electron track
- j) rms deviation of electron hits from the fitted track
- k) number of tracks identified as hadrons in the event

Efficiency Typically 25-40%:  
Algorithms under development

$$\text{FOM} = \text{Signal} / \text{Sqrt}(\text{Bkg})$$

| Cell width | Cell Depth | Relative FOM | Electron Energy |
|------------|------------|--------------|-----------------|
| 3.8        | 4.5        | 1.0          | 10.0%           |
| 3.8        | 6.0        | 1.02         | 10.7%           |
| 5.4        | 4.5        | 0.9          | 9.9%            |





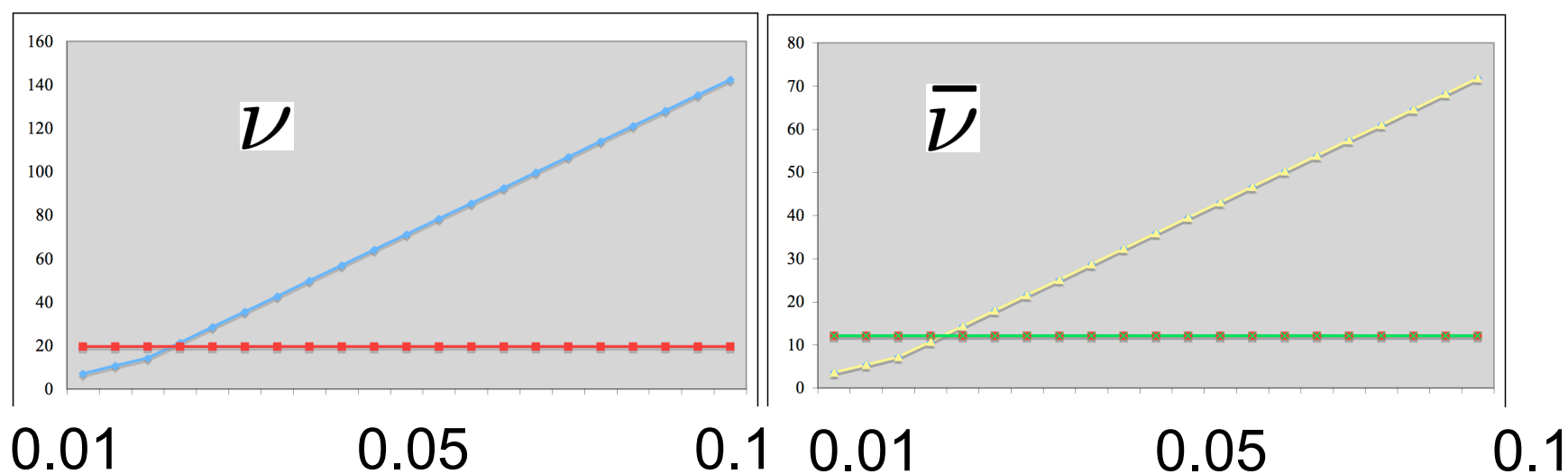
# Tradeoffs

- Primary problem is tradeoff between acceptance and background rejection:
  - E.g.  $\pi^0 \rightarrow \gamma\gamma \rightarrow \text{electrons}$
  - Look for “gap” from photons at beginning, followed by shower
  - Tradeoff of acceptance vs. shower fluctuations vs. background rate and ability to accurately model
  - Beam-Related backgrounds are measurable and small
- Lots to do, progress being made



# Signal and Backgrounds

- Statistical Power: why this is hard and we need protons



For  $\sin^2 2\theta_{13} = 0.1$ :

$\nu$ : S=142.1, B=19.5

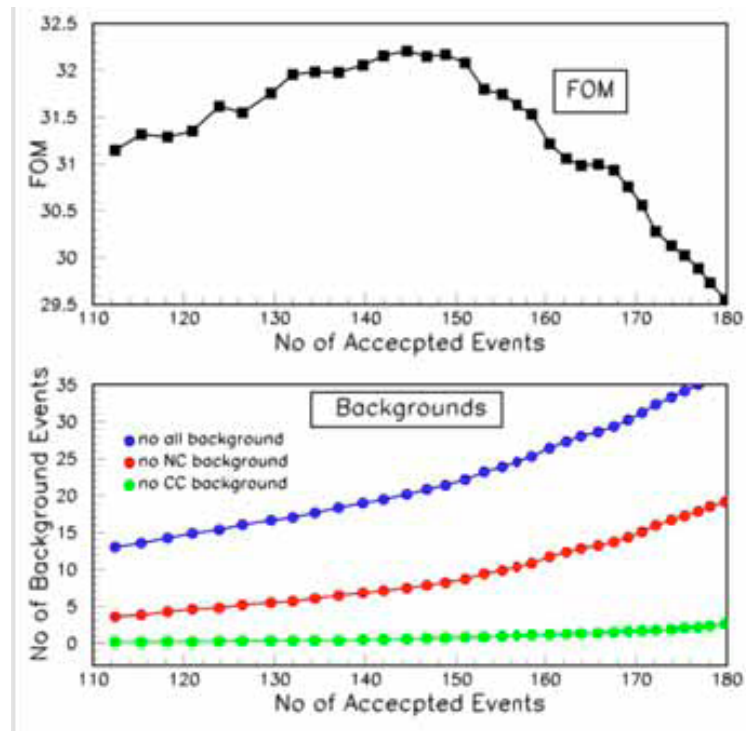
$\bar{\nu}$ : S= 71.8, B=12.1

5 yrs at 6.5E20 pot/yr,  
efficiencies included



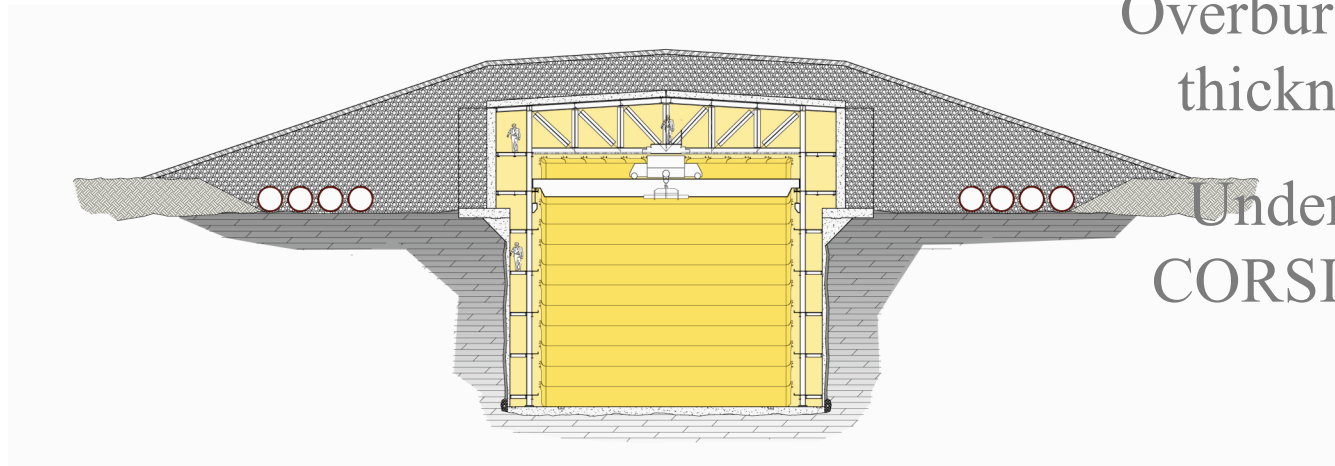
# Sensitivity to Reconstruction Cuts

- Vary cut on likelihood function
- What Happens to  $\text{Signal}/\sqrt{\text{Bkg}} = \text{FOM}$
- What happens to rates?
- *Nothing drastic*
- *Very much under development, but we're thinking about this and the numbers are reasonable*



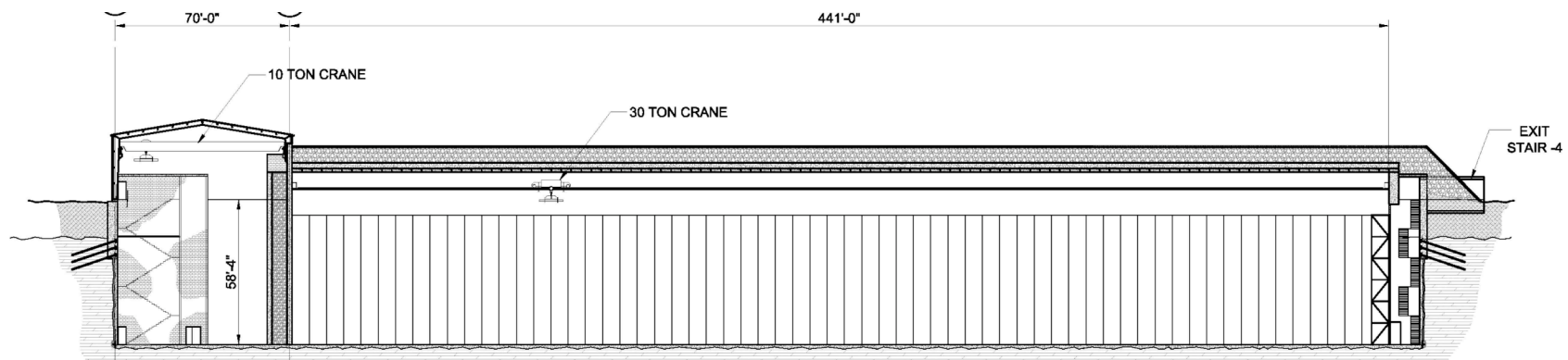


# Far Detector Building Design with Overburden



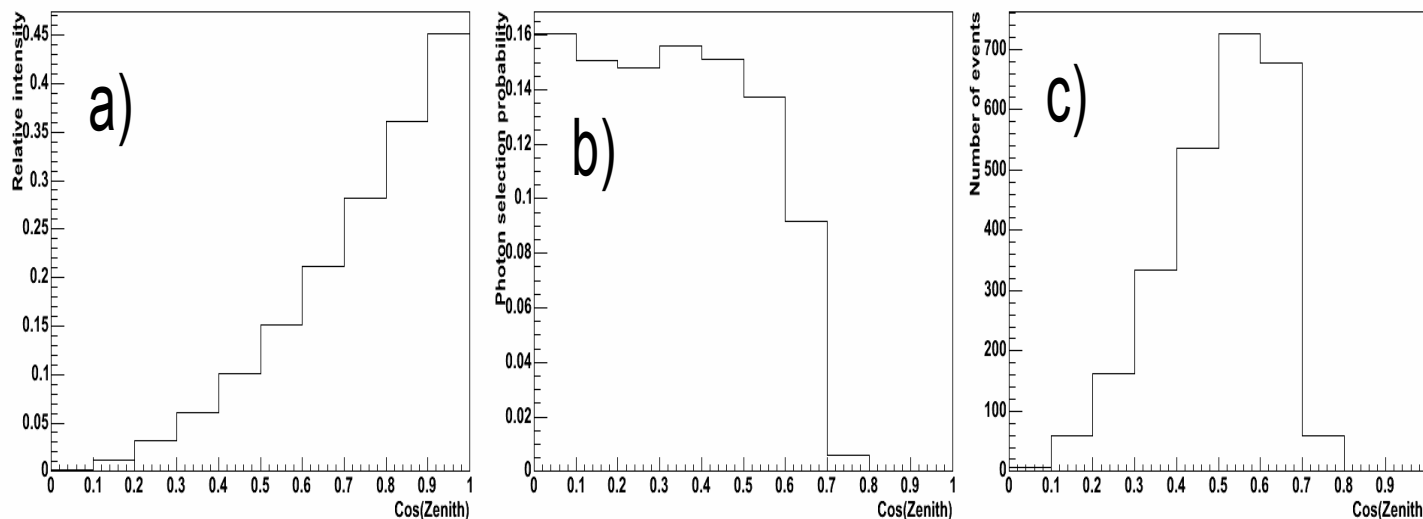
Overburden: Minimum  
thickness possible

Under study with  
CORSIKA/GEANT





# Determination of Overburden



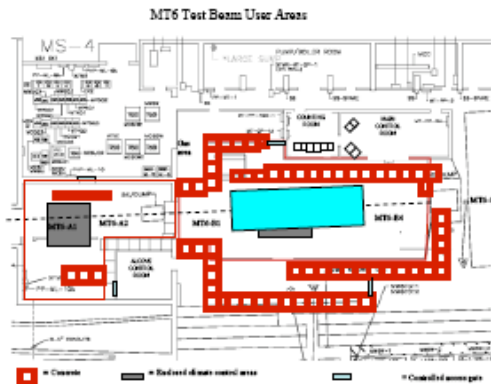
- a) Incident Rate
- b) Probability
- c) Convolution

$\gamma$ -flux is source of bkg:  
needs to come in at angle  
to look like signal

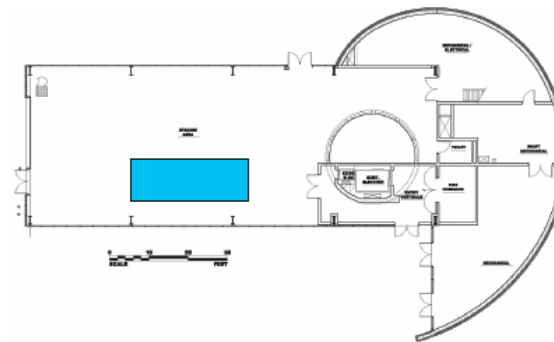
Need 2 m overburden to eliminate



# Near Detector: Modular & Mobile



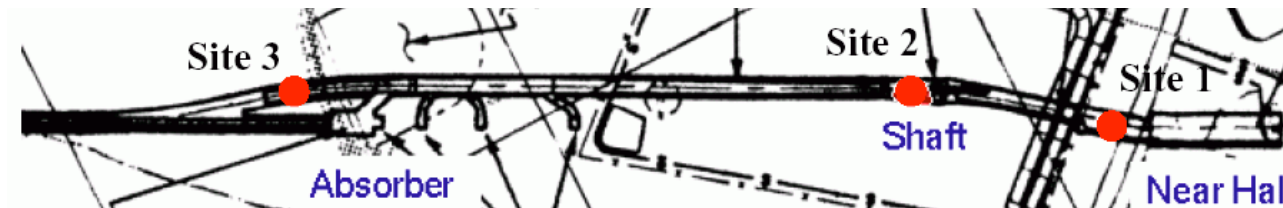
M Test



MINOS Surface Building

*Response to  
 $\nu_\mu$  and  $\nu_e$*

*Calibrate in Test Beam*

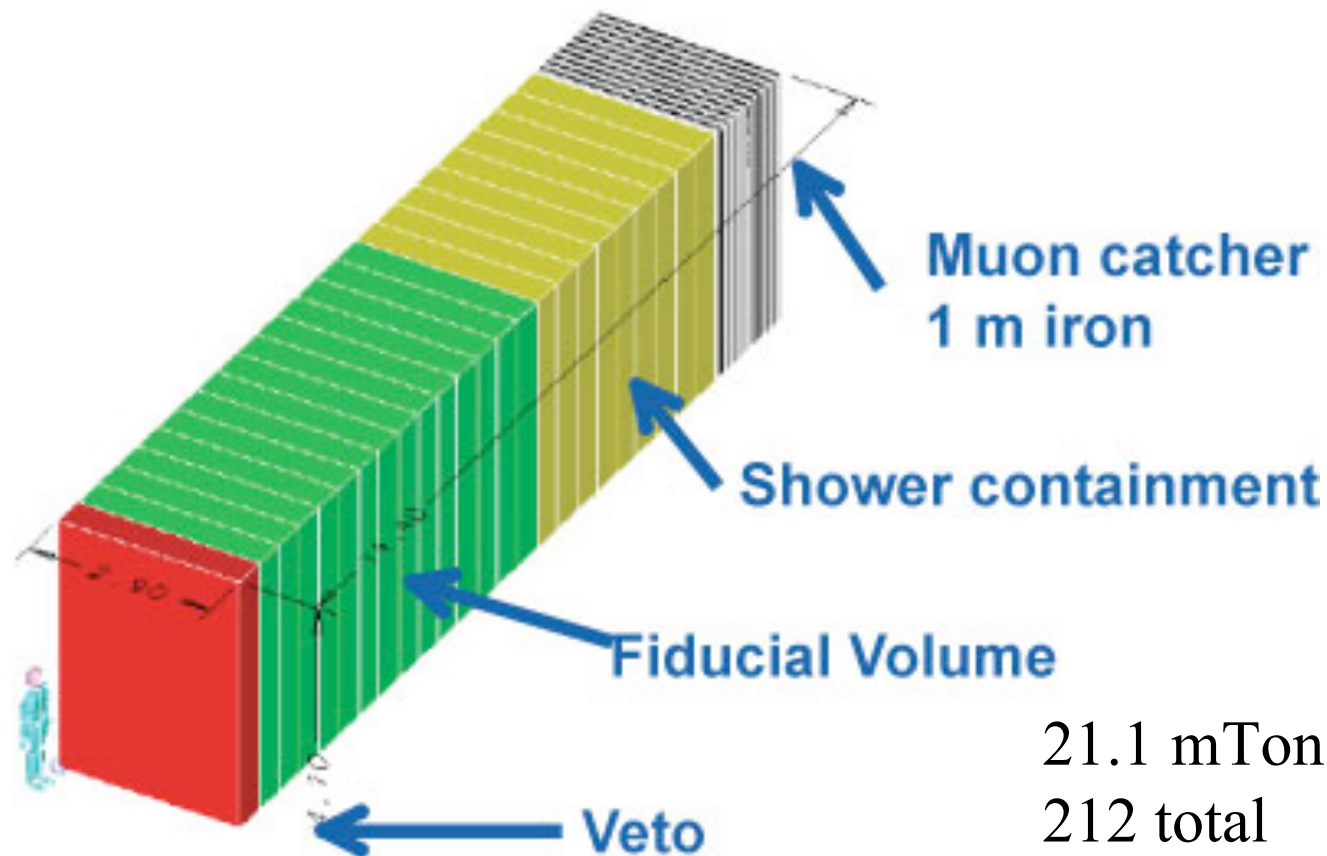


NuMI Access Tunnel

*Measure Beam  
Properties*



# Near Detector

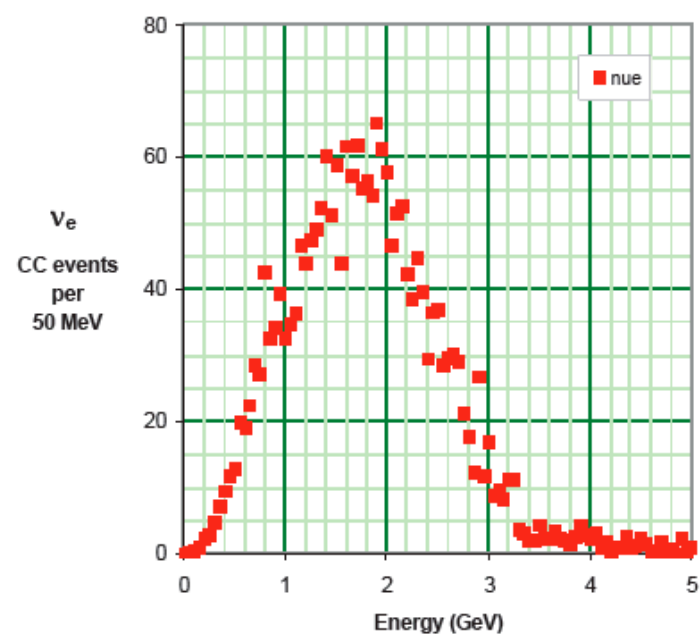
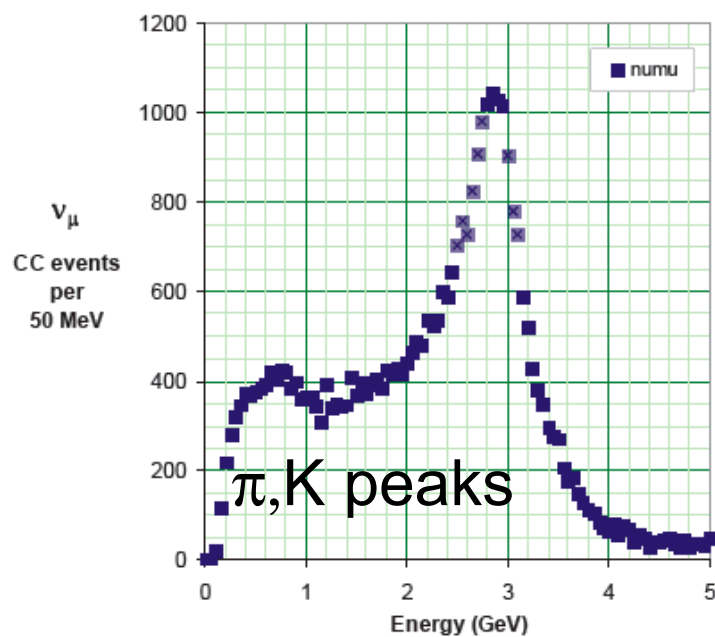


21.1 mTons fiducial  
212 total  
3.2 events/pulse



# Surface Measurement

$6.5 \times 10^{20}$  pot in 75 mrad off-axis beam



Far Data Spectrum peaked at 2 GeV





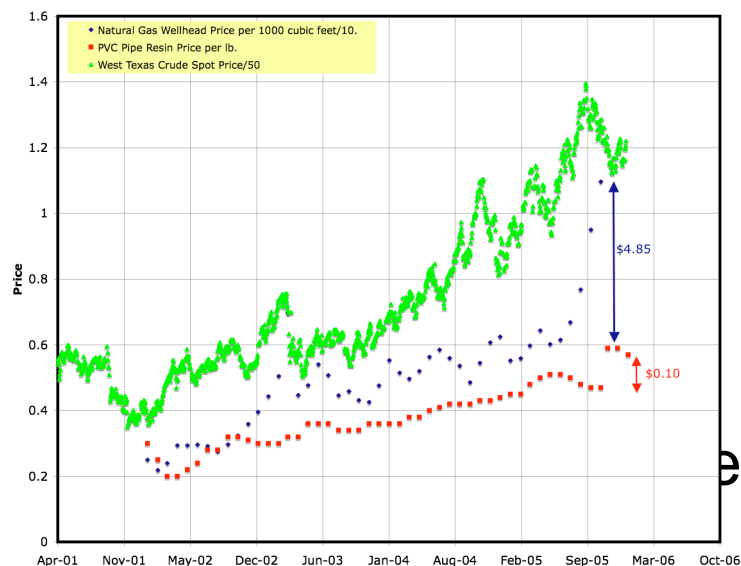
# Cost

|                         | Contingency | Total Cost M\$ |
|-------------------------|-------------|----------------|
| Far Detector            |             |                |
| Active detector         | 30%         | 80             |
| Electronics and DAQ     | 55%         | 13             |
| Shipping                | 21%         | 7              |
| Installation            | 43%         | 14             |
| Near Detector           | 44%         | 3              |
| Building and outfitting | 58%         | 29             |
| Project management      | 25%         | 5              |
| Additional contingency  |             | 14             |
| Total                   | 50%         | 165            |

- Sensitive to price of crude oil!
  - We have understanding of how price depends on petroleum price: can make sensible estimates

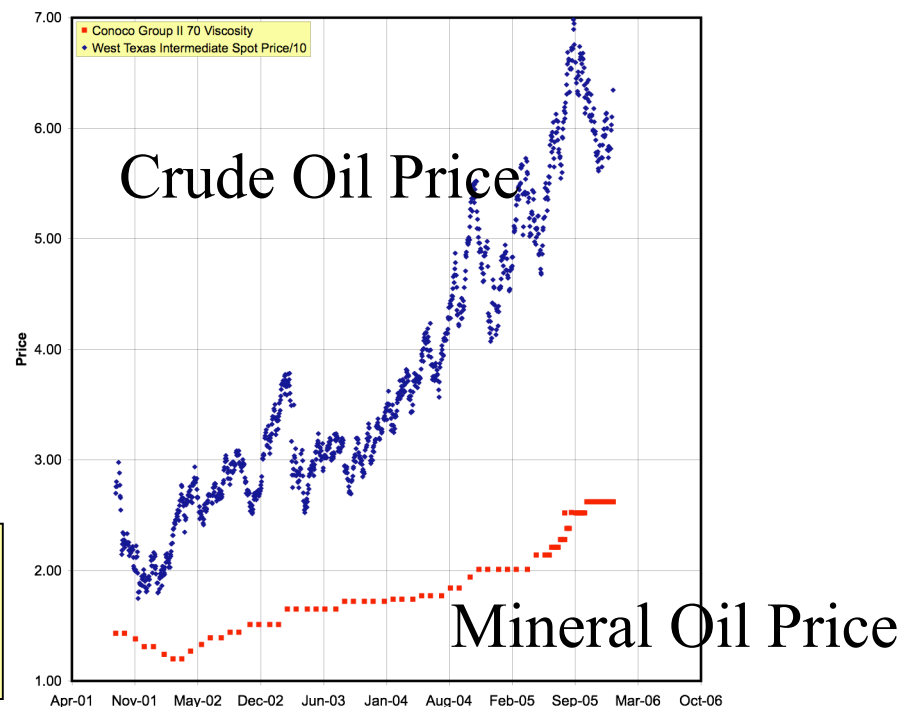


# Sensitivity to Petroleum Price



PVC resin increases ~ \$0.03/lb.  
for every \$10/barrel increase in  
the cost of crude oil.

Mineral oil increases ~\$0.03/gal for  
every \$1.00 increase in the cost of  
crude.





# Proton Intensity

- NOvA will run after the Tevatron terminates operation. Thus, parts of the accelerator complex now devoted to antiproton production and storage will be available for the neutrino program.
- Assumption of our proposal:
  - 0.7 MW, of which we assumed 0.625 MW for the NOvA program. (for no proton driver)
- Fermilab strategy:
  - If ILC looks affordable, move to host the ILC and do “cheap” upgrades to the proton intensity.
  - If ILC will be delayed, move toward a Proton Driver (i.e., a new Booster).
- We assume that the Proton Driver will allow 2.4 MW.



# Schedule

- Cannot start in FY2007 due to regulations on Congressional line items.
- Aiming for a FY2008 project start.
- Start of data taking October 2010
- Completion of the Far Detector July 2011

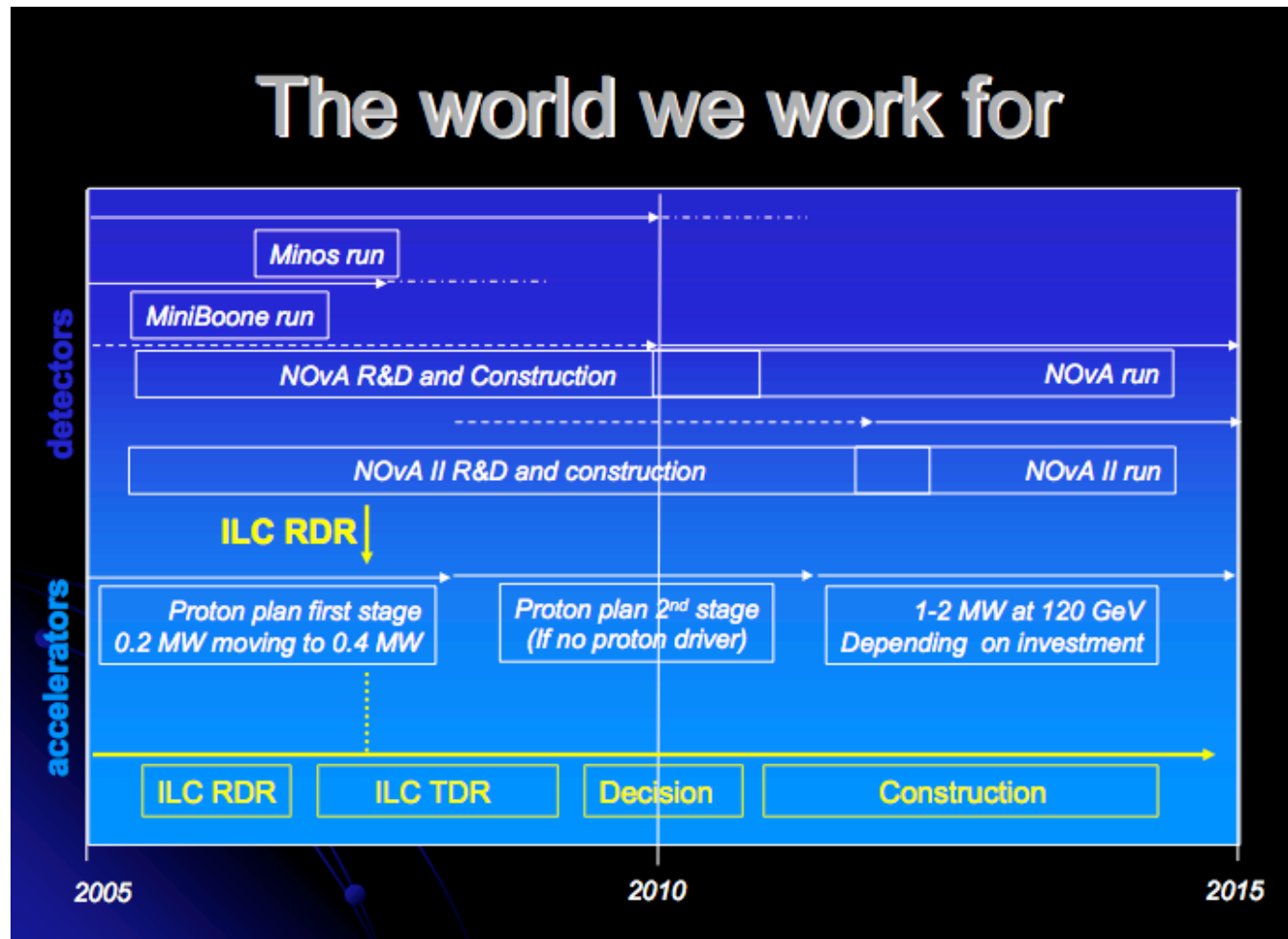
3 Feb 06: President's Budget: White House to Congress  
High Energy Physics Program (\$775.1 million)  
This is a \$58.4 million increase over FY 2006.



*Project engineering and design funding of \$10.3 million is requested for the new Electron Neutrino Appearance project.*



# From Oddone to EPP2010





# Conclusions

- The Fermilab/NuMI/NOvA program provides a flexible, step-by-step approach:

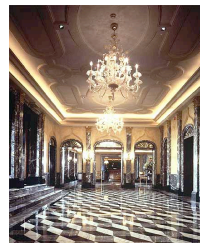
Gateway  
to  
Paradise:  
 $\sin^2 2\theta_{13}$



Ghiberti's Gates of Paradise in Firenze



Mass  
Hierarchy  
(leaving, not  
coming in!)



CP  $\delta$

$\Delta m^2_{23}, \theta_{23}$   $\mathcal{O}(2800 \text{ kg Au in mosaics})$